

**Goal:** Make sense of Wasserstein distances in high dimension by designing a robust variant of the Wasserstein.

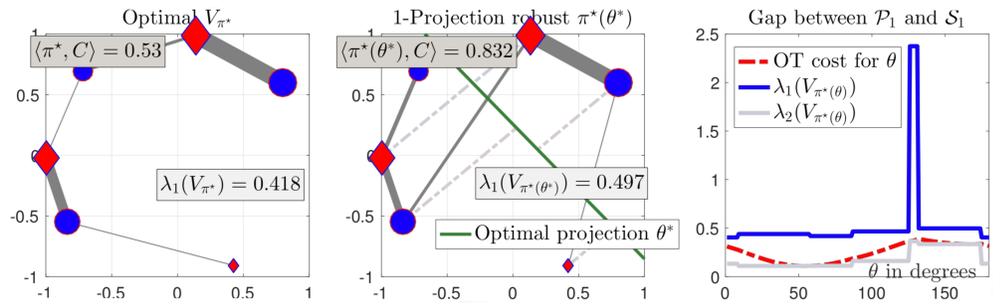
**Approach:** Project the measures onto a low-dimensional subspace and consider the maximum over all subspaces.

**Results:** Geodesic metric equivalent to  $\mathcal{W}$ . Efficient algorithms and use case on text data.

## I. Wasserstein Distance in High Dimension

$$\mathcal{W}^2(\mu, \nu) = \inf_{\pi \in \Pi(\mu, \nu)} \int \|x - y\|^2 d\pi(x, y)$$

- Problems:**
- $\|\cdot\|^2$  not informative in high dimension
  - $|W_2^2(\hat{\mu}, \hat{\nu}) - W_2^2(\mu, \nu)| \sim \left(\frac{1}{n}\right)^{1/d}$



## II. Projection and Subspace Robust Wasserstein Distances

### Projection Robust Wasserstein Distance (PRW)

$$\mathcal{P}_k(\mu, \nu) = \sup_{\dim(E)=k} \mathcal{W}(P_E \# \mu, P_E \# \nu)$$

Not convex !

### Subspace Robust Wasserstein Distance (SRW)

- Corresponding "min-max" problem:

$$\mathcal{S}_k^2(\mu, \nu) = \inf_{\pi \in \Pi(\mu, \nu)} \sup_{\dim(E)=k} \int \|P_E(x) - P_E(y)\|^2 d\pi(x, y)$$

- SRW is a convex relaxation of PRW:

$$\mathcal{S}_k(\mu, \nu) = \max_{\substack{0 \leq \Omega \leq I \\ \text{trace}(\Omega)=k}} \mathcal{W}(\Omega^{1/2} \# \mu, \Omega^{1/2} \# \nu)$$

- SRW finds a coupling  $\pi$  minimizing the spectral cost:

$$\mathcal{S}_k^2(\mu, \nu) = \min_{\pi \in \Pi(\mu, \nu)} \sum_{l=1}^k \lambda_l(V_\pi)$$

Where  $V_\pi$  is the Second Order Moment Matrix of the Displacements:

$$V_\pi = \int (x - y)(x - y)^T d\pi(x, y)$$

## IV. Computing SRW

### Entropic Regularization

- Ensures uniqueness of optimal  $\pi^*$
- Sinkhorn algorithm

### Frank-Wolfe

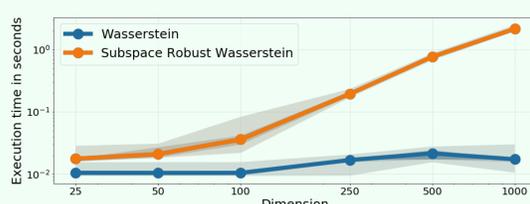
**Algorithm** Frank-Wolfe algorithm for entropic SRW  
**Input:** Measures  $(x_i, a_i)$  and  $(y_j, b_j)$ , dimension  $k$ , regularization strength  $\gamma > 0$   
 Initialize  $\Omega$   
**for**  $t = 0$  to  $\text{max\_iter}$  **do**  
 $\pi \leftarrow \text{reg\_OT}((x, a), (y, b), \text{reg} = \gamma, \text{cost} = d_\Omega^2)$   
 $U \leftarrow \text{top } k \text{ eigenvectors of } V_\pi$   
 $\tau = 2 / (2 + t)$   
 $\Omega \leftarrow (1 - \tau)\Omega + \tau [U \text{diag}([1_k, 0_{d-k}]) U^T]$   
**end for**  
**Output:**  $\Omega, \pi, \langle \Omega | V_\pi \rangle$

### Projected Gradient Method

**Algorithm** Projected supergradient method for SRW  
**Input:** Measures  $(x_i, a_i)$  and  $(y_j, b_j)$ , dimension  $k$   
 Initialize  $\Omega$   
**for**  $t = 0$  to  $\text{max\_iter}$  **do**  
 $\pi \leftarrow \text{OT}((x, a), (y, b), \text{cost} = d_\Omega^2)$   
 $\Omega \leftarrow \text{Proj}[\Omega + \frac{1}{t+1} V_\pi]$   
**end for**  
**Output:**  $\Omega, \langle \Omega | V_\pi \rangle$

### Computation Time

- Warm start in Sinkhorn
- Quadratic in dimension  $d$



## III. The SRW Geometry

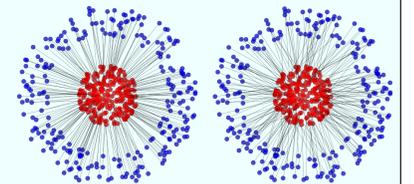
SRW is a distance between probability measures

SRW is equivalent to the 2-Wasserstein distance

$$\sqrt{\frac{k}{d}} \mathcal{W}(\mu, \nu) \leq \mathcal{S}_k(\mu, \nu) \leq \mathcal{W}(\mu, \nu)$$

### Geodesics in SRW space

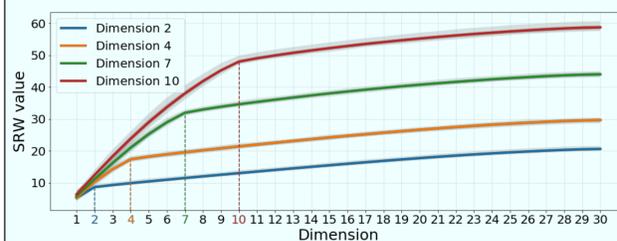
$\pi^* \in \Pi(\mu, \nu)$  minimizing  $\pi \mapsto \sum_{l=1}^k \lambda_l(V_\pi)$



SRW (left) and  $\mathcal{W}$  (right) geodesics in presence of noise ( $d=30$ )

### Dependence on dimension

$k \mapsto \mathcal{S}_k^2(\mu, \nu)$  increasing and concave

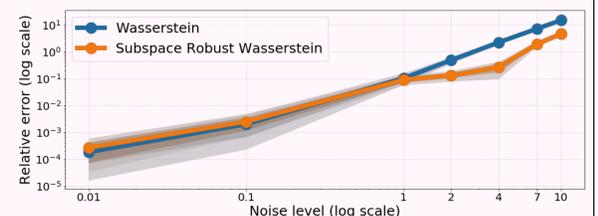


- Measures in dimension  $d=30$ , with transport only occurring in dimension 2, 4, 7 and 10 respectively.
- Use a 'elbow' rule of thumb to choose  $k$  in practice.

## V. Applications

### SRW is Robust to Noise

Low-dimensional Gaussians are added noise. We plot the relative error for SRW and  $\mathcal{W}$  distances.



### Movies

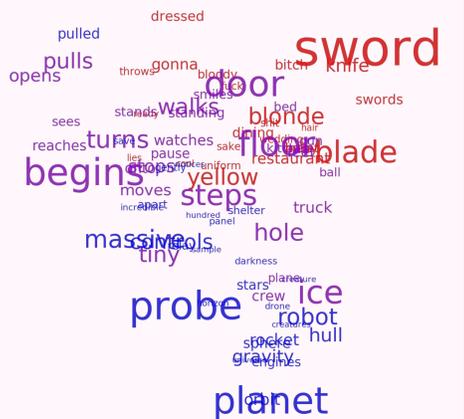
Scenarios of movies are transformed into measures in  $\mathbb{R}^{300}$  using Word2vec

KILL BILL VOLUME 2  
KILL BILL VOLUME 1

TITANIC  
DUNKIRK

THE MARTIAN  
INTERSTELLAR  
GRAVITY

Metric MDS of SRW distances between movies



Optimal 2D-subspace between Kill Bill (red) and Interstellar (blue)